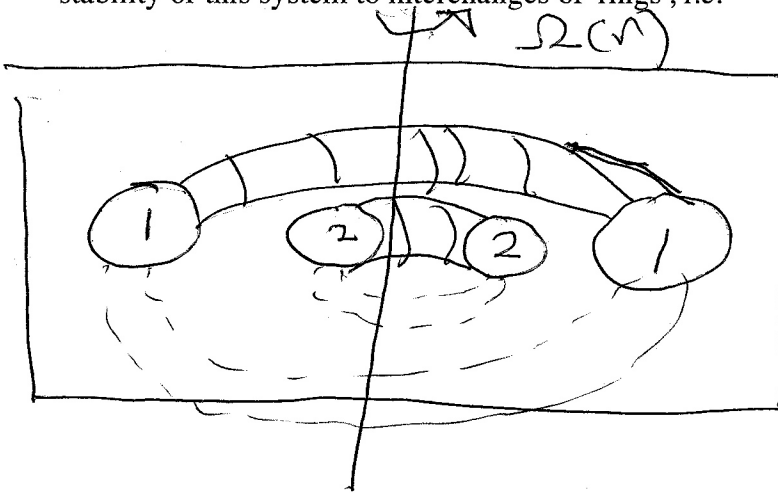


Problem Set V: Due Thursday, March 15.

- 1.) Consider a fluid in hydrostatic equilibrium with a vertical entropy gradient $\partial S/\partial z < 0$. Take $\underline{g} = -g\hat{z}$.
 - a.) Starting from the basic equations, derive the growth rate of ideal Rayleigh-Bernard instability. You will find it helpful to relate the density perturbation $\tilde{\rho}/\rho_0$ to the temperature perturbation \tilde{T}/T_0 by exploiting the fact that the instability develops slowly in comparison to the sound transit time across a cell. Relate your result to the Schwarzschild criterion discussed in class.
 - b.) Now, include thermal diffusivity (χ) and viscosity (ν) in your analysis. Calculate the critical temperature gradient for instability, assuming $\chi \sim \nu$. Discuss how this compares to the ideal limit. What happens if $\nu > \chi$?

- 2.) Now again, consider the system of Problem 2, now immersed in a uniform magnetic field $\underline{B} = B_0\hat{z}$.
 - a.) Assuming ideal dynamics, use the Energy Principle to analyze the stability of a convection cell of vertical wavelength k_z . Of course, $k_z L_p \gg 1$, where L_p is a mean pressure scale length. What is the effect of the magnetic field? Can you estimate how the growth rate changes?
 - b.) Now, calculate the growth rate using the full MHD equations. You may assume $\underline{\nabla} \cdot \underline{V} = 0$. What structure convection cell is optimal for vertical transport of heat when B_0 is strong? Explain why. What happens when $B_0 \rightarrow \infty$? Congratulations - you have just derived a variant of the Taylor-Proudman theorem!

- 3.) Consider a rotating fluid with mean $\underline{V} = r\Omega(r)\hat{\theta}$. Your task here is to analyze the stability of this system to interchanges of 'rings', i.e.



In all cases, assume $\underline{\nabla} \cdot \underline{V} = 0$ and $k_\theta = 0$, so the interchange motions carry no angular momentum themselves and the cells sit in the r - z plane.

- a.) At the level of a "back-of-an-envelope" calculation, calculate the change in energy resulting from the incompressible interchange of rings (1) and (2). Note that $E = L^2/2mr^2$ and that the angular momentum L of an interchanged ring is conserved, since $k_\theta = 0$. From this, what can you conclude about the profile of $\Omega(r)$ necessary for stability? Congratulations - you have just derived the Rayleigh criterion!
- b.) Now, calculate the interchange growth rate by a direct solution of the fluid equations. You may find it helpful to note that for rotating fluids in cylindrical geometry:

$$\frac{\partial V_r}{\partial t} + \underline{V} \cdot \underline{\nabla} V_r - \frac{V_\theta^2}{r} = \frac{-1}{\rho} \frac{\partial P}{\partial r},$$

$$\frac{\partial V_\theta}{\partial t} + \underline{V} \cdot \underline{\nabla} V_\theta + \frac{V_r V_\theta}{r} = \frac{-1}{\rho r} \frac{\partial P}{\partial \theta},$$

$$\frac{\partial V_z}{\partial t} + \underline{V} \cdot \underline{\nabla} V_z = \frac{-1}{\rho} \frac{\partial P}{\partial z}.$$

Show that you recover the result of part (a).

- c.) Compare and contrast this interchange instability to an incompressible Rayleigh-Taylor instability. Make a table showing the detailed correspondences.
- 4.) This problem asks you to explore the Current Convective Instability (CCI) in a homogeneous medium and its sheared field relative, the Rippling Instability.
- a.) Consider first a current carrying plasma in a straight magnetic field $\underline{B} = B_0 \hat{z}$ - i.e. ignore the poloidal field, etc. Noting that the resistivity η is a function of temperature (ala' Spitzer - c.f. Kulsrud 8.7), calculate the electrostatic resistive instability growth rate, assuming T evolves according to:

$$\frac{\partial T}{\partial t} + \underline{v} \cdot \nabla T - \chi_{\parallel} \partial_z^2 T - \chi_{\perp} \nabla_{\perp}^2 T = 0$$

and the electrostatic Ohm's Law is just

$$-B_0 \partial_z \phi = \frac{1}{\eta} \frac{d\eta}{dt} \tilde{T}(\eta J_0).$$

- b.) *Thoroughly* discuss the physics of this simple instability, i.e.
- what is the free energy source?
 - what is the mechanism?
 - what are the dampings and how do they restrict the unstable spectrum?
 - how does spectral asymmetry enter?
 - what is the cell structure?
- c.) Use quasilinear theory and the wave breaking limit to estimate the heat flux from the C.C.I.

- d.) Now, consider the instability in a *sheared* magnetic field.
- i.) What difficulties enter the analysis?
 - ii.) Resolve the difficulty by considering coupled evolution of vorticity, Ohm's Law (in electrostatic limit but with temperature fluctuations) and electron temperature. Compute the growth rate in the limit $\chi_{\parallel}, \chi_{\perp} \rightarrow 0$. Compute the mode width. Discuss how asymmetry enters here. Explain why.
- e.) Noting that $\chi_{\parallel} \gg \chi_{\perp}$ (why? - see Kulsrud 8.7), estimate when parallel thermal conduction becomes an important damping effect. Can χ_{\parallel} alone ever absolutely stabilize the rippling mode?
- f.) Calculate the quasilinear heat flux and use the breaking limit to estimate its magnitude.
- 5a.) Derive the quasilinear equation for the evolution of the pressure profile in a plasma which supports resistive interchange turbulence. Express your answer in terms of the radial displacement spectrum $\left| \tilde{\epsilon}_{rk} \right|^2$.
- b.) Assuming localized heating with on-axis central deposition, derive an expression for the stationary pressure gradient in terms of the intensity profile of the resistive interchange displacement spectrum.
- c.) Using the wave-breaking estimate for saturation level ($\tilde{\epsilon}_r \sim k_r^{-1}$), calculate how β scales with input power, assuming fixed density. When does this scaling fail? Why?
- 6.) Read K.V. Roberts and J.B. Taylor, *Physics of Fluids*, Vol. 8, pg 315, 1965. Write a short essay (including cartoons and schematic equations) explaining the "point" of this paper. Contrast the "twisted slicing modes" it describes with the resistive interchange mode discussed in class.